

# *SCRAAM – AN EXPERIMENT TO TEST REACTOR FLUX PREDICTIONS AND PROBE THE "REACTOR ANTINEUTRINO ANOMALY"*

November 17, 2011



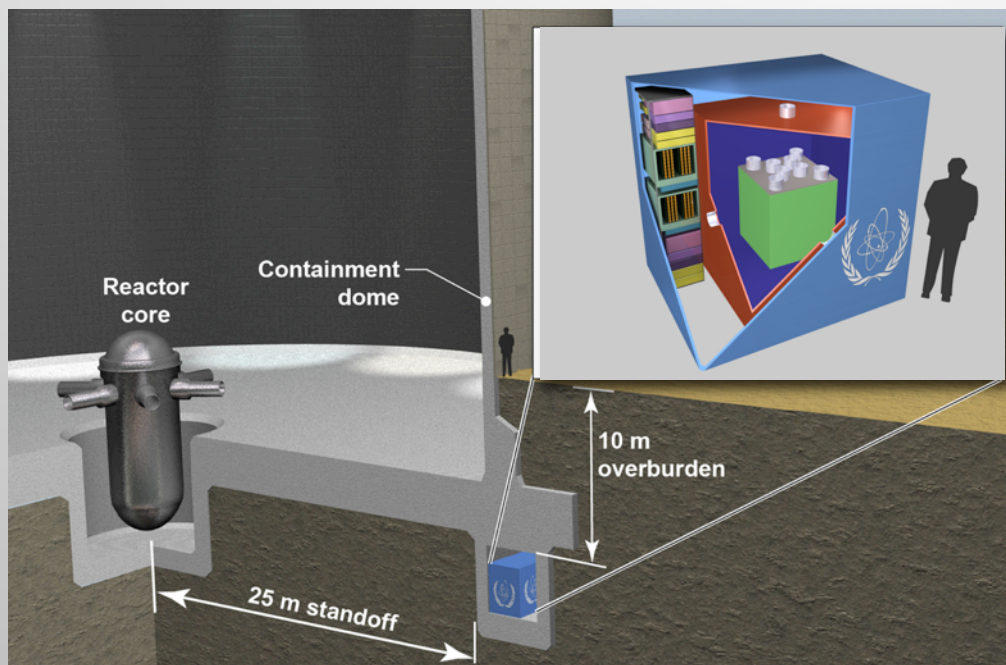
Nathaniel Bowden

LLNL-PRES-514073

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# There is increasing interest in (Short Baseline) Antineutrino Monitoring of Reactors



## AGENDA

Ad Hoc Working Group on Safeguards Applications of Antineutrino Detectors, 14 September 2011, Vienna, Austria



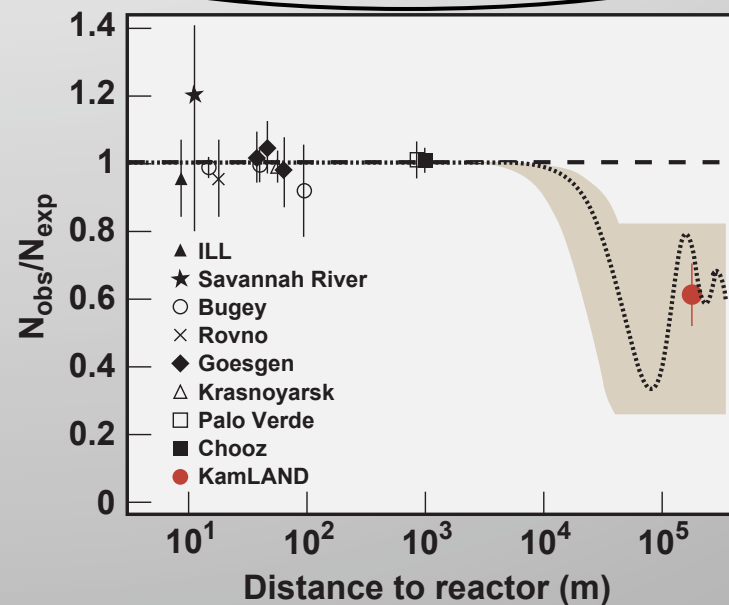
# Basic science laid the foundation for this monitoring technique

- Reines and Cowan, 1956:
  - First to detect antineutrinos using a reactor source and a liquid scintillator detector



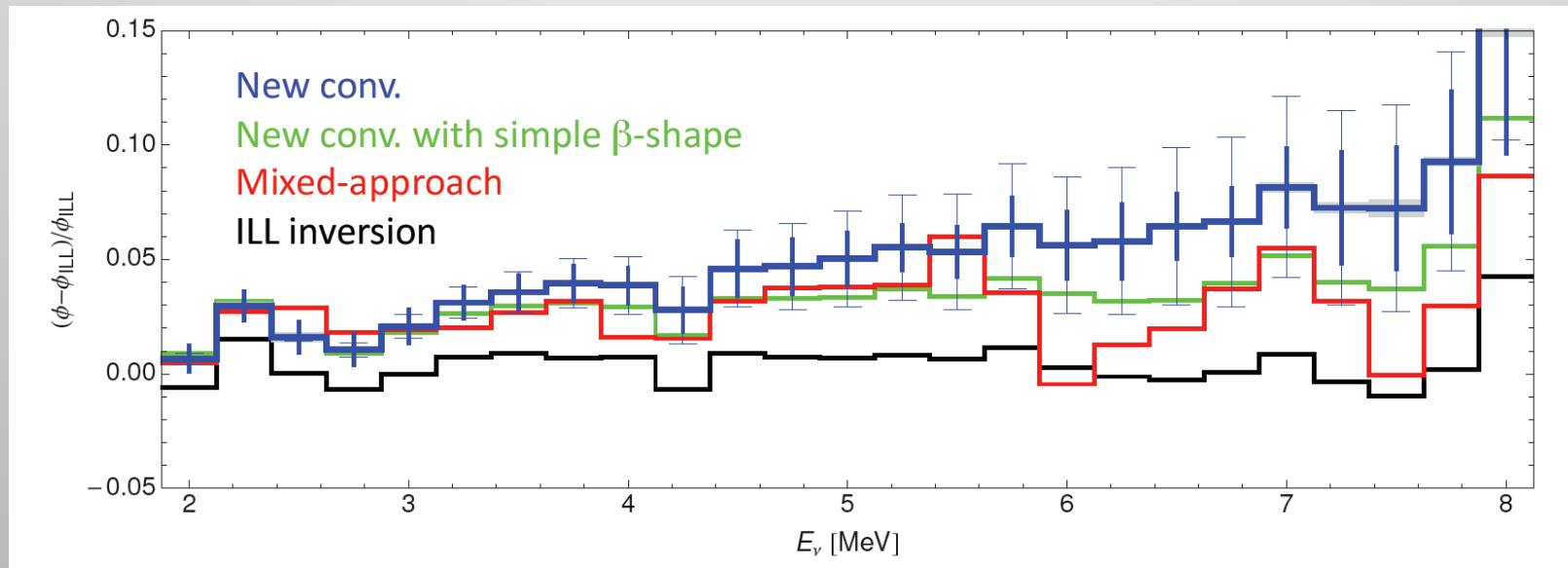
- Three decades of neutrino oscillation studies have provided:
  - A mature technology base
  - A quantitative understanding of reactors as an antineutrino source

?



# Recent Re-evaluations of the Reactor Antineutrino Flux

- We have seen strong efforts to improve the conversion of ILL reference spectra to antineutrino spectra:  $\sim 3\%$  increase in flux
- Two largely independent predictions agree:



Mueller, et al: arXiv:1101.2663  
Huber: arXiv:1106.0687

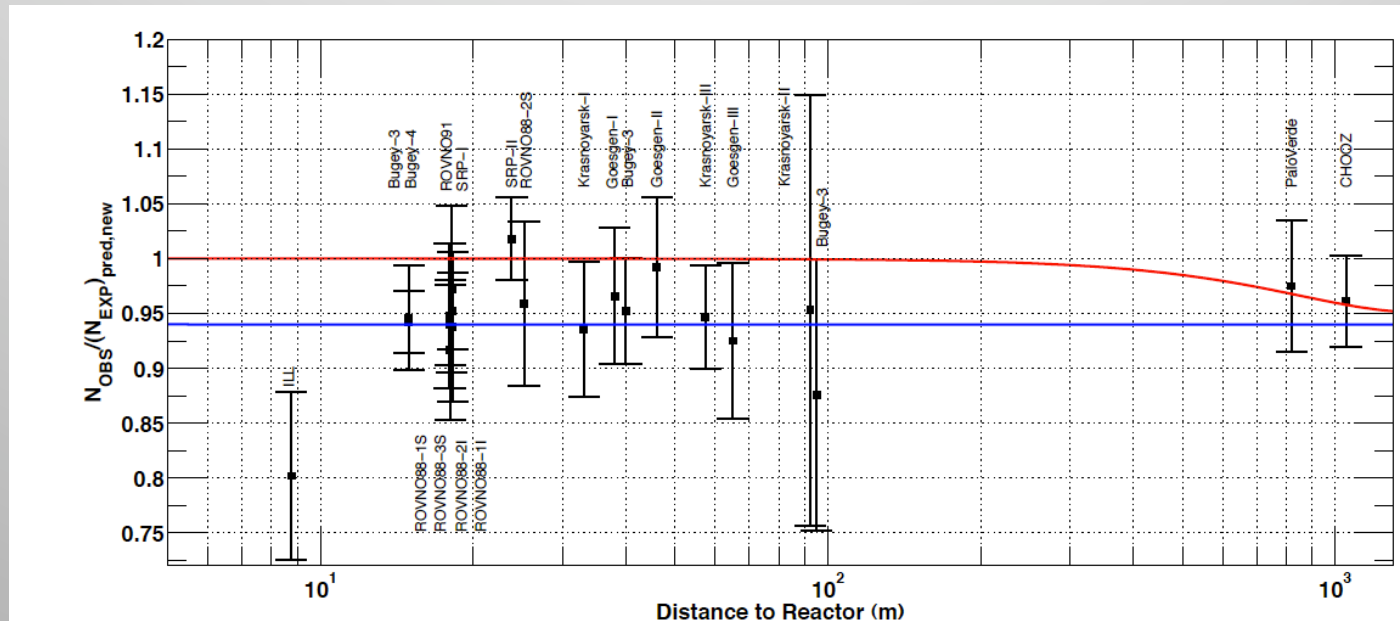
- But, there are still considerable uncertainties related to some corrections:
  - a high-precision spectral measurement might help



# The Reactor Antineutrino Anomaly

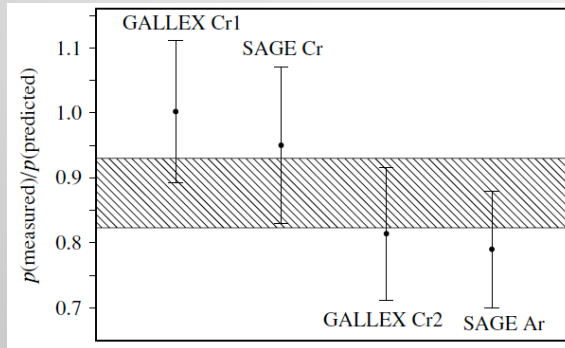
- Mention, et al, re-analyzed many previous short baseline reactor experiments, in light of their new antineutrino flux prediction
- The result: new global “Reactor Antineutrino Anomaly”

$$N_{\text{obs}}/N_{\text{pred}} = 0.979 \pm 0.029 \Rightarrow 0.943 \pm 0.023$$

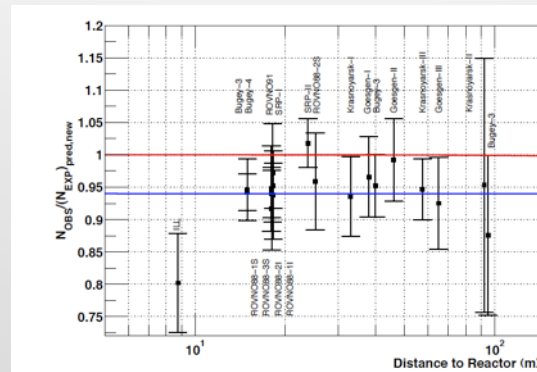


arXiv:1101.2755v4

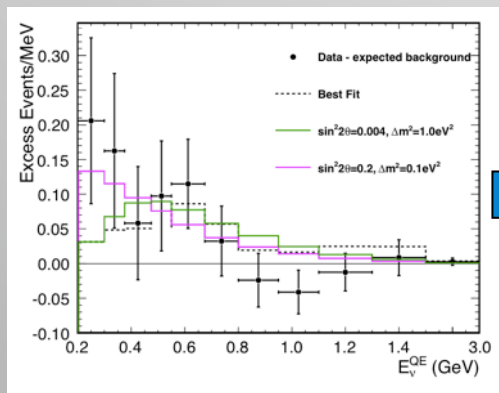
# The Reactor Anomaly is consistent with other hints at a sterile flavor



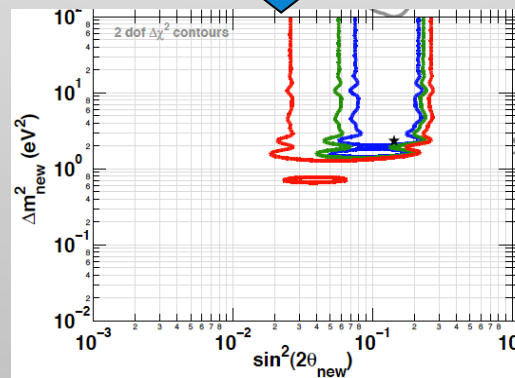
SAGE/GALLEX



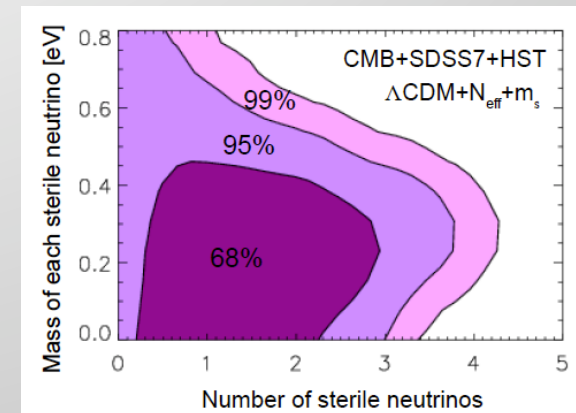
Reactor Anomaly



MiniBOONE



Combined



Astrophysical measurements are also consistent with  $\sim \text{eV}$  sterile(s)

# The recent results have sparked a new flurry of interest and activity

**STERILE NEUTRINOS AT THE CROSSROADS**  
September 25-28, 2011 • Blacksburg, VA • USA

**Short-Baseline  
Neutrino Workshop**  
*12-14 May 2011*

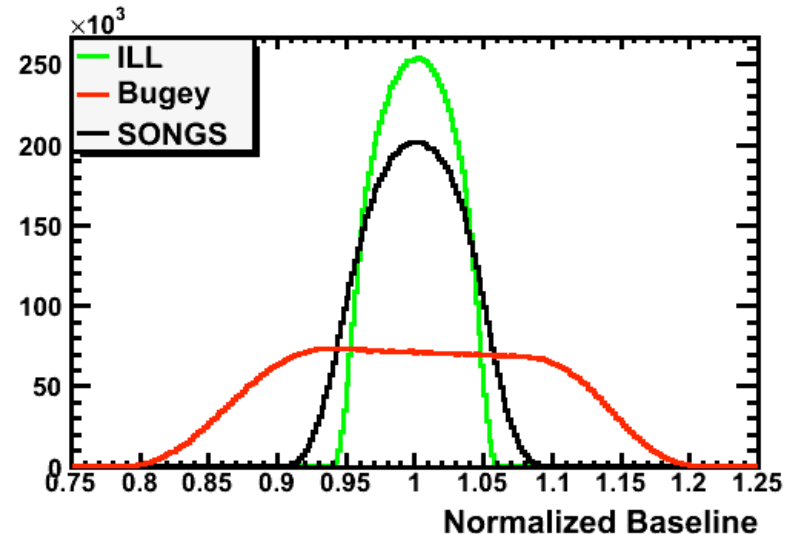
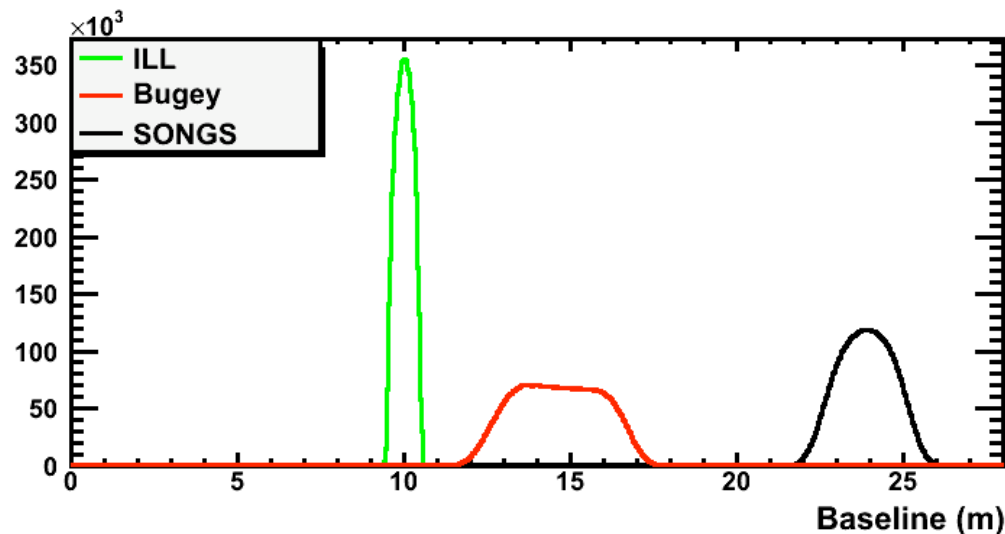


- Most sterile hints take the form of a deficit or excess relative to an (uncertain) expectation
- Strong desire in community for definitive experiments based on measurement of oscillation patterns
- Can a new short baseline reactor experiment help?

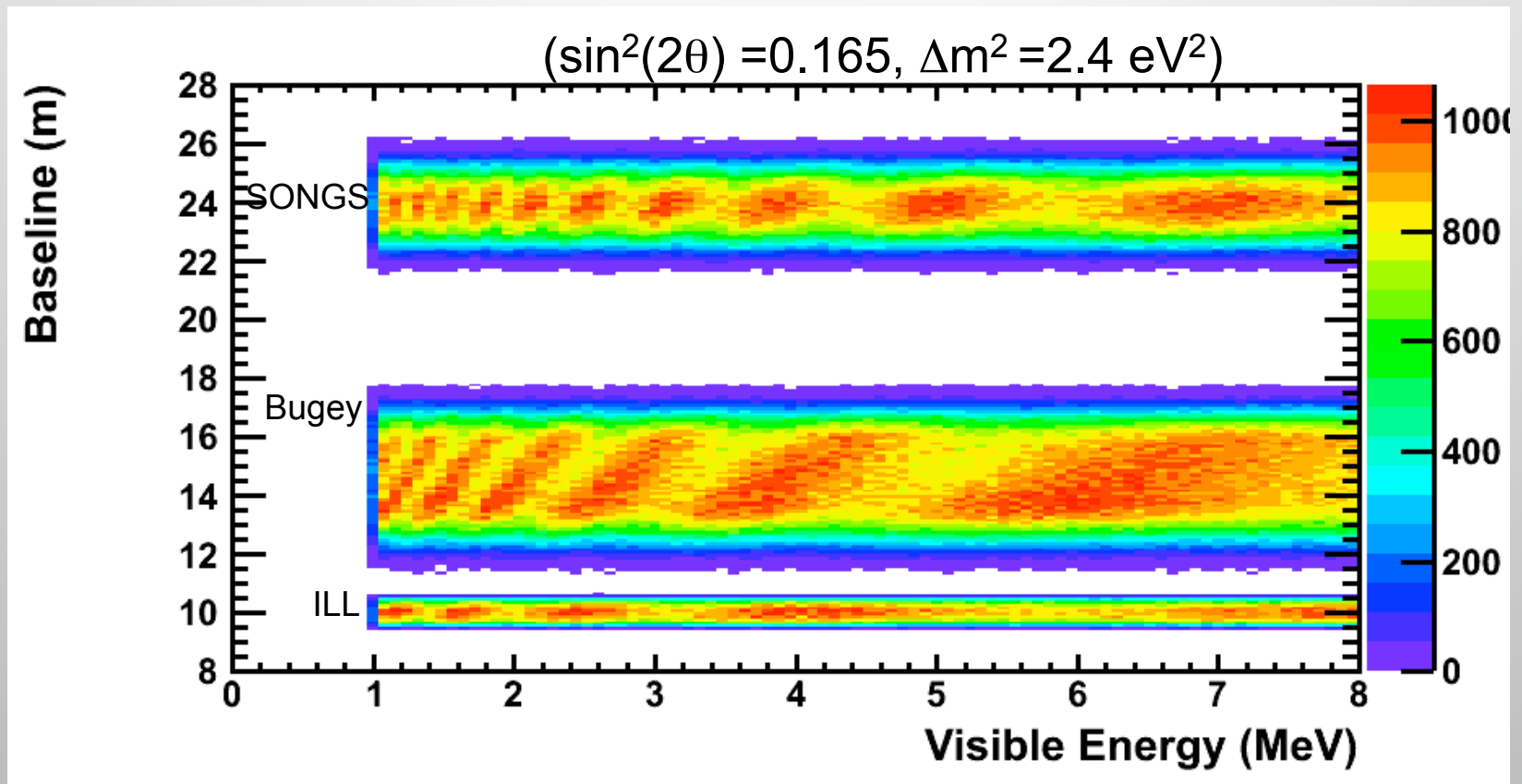


# At short baselines, a reactor is not a point source

Reactor	Baseline	Core	Detector	$\Delta L/L$ (FWHM)	Power	Flux $\nu/m^2/s$
ILL	10m	$\varnothing 0.4m \times 0.2m$ (HEU)	$\varnothing 1m \times 1m$	$\sim 8\%$	$58 \text{ MW}_{th}$	$\sim 1 \times 10^{16}$
Bugey3	15m	$\varnothing 2.5 \times 2.5m$	$1m \times 1m$	$\sim 30\%$	$2800 \text{ MW}_{th}$	$\sim 2 \times 10^{17}$
SONGS	24m	$\varnothing 3m \times 2m$	$\varnothing 1m \times 2m$	$\sim 10\%$	$3400 \text{ MW}_{th}$	$\sim 1 \times 10^{17}$



# Effect of Baseline and Baseline Distribution



- No previous experiment appears to have been optimized in this respect
- Experiments at appropriate small and large reactors would be complementary:
  - efficiently probe different  $\Delta m^2$  regions
  - measuring flux/spectra from different core compositions

# SCRAAM: The Southern California Reactor Antineutrino Anomaly Monitor

- Our proposal is to perform a relatively rapid and inexpensive experimental measurement
  - Direct sterile oscillation sensitivity via spectra distortion
  - High statistics flux and spectrum measurement from a single Pressurized Water Reactor (PWR)
- This requires access to location(s) with high antineutrino flux and appropriate core-detector geometry

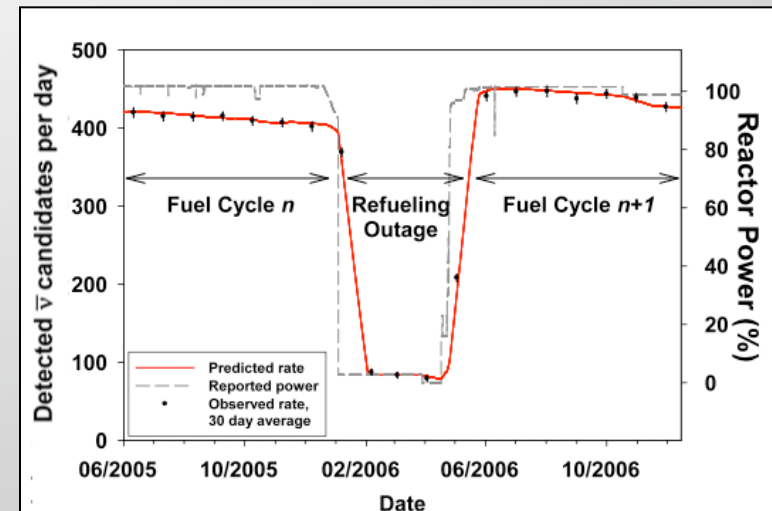




# The San Onofre Nuclear Generating Station: Our (nonproliferation) laboratory for over a decade

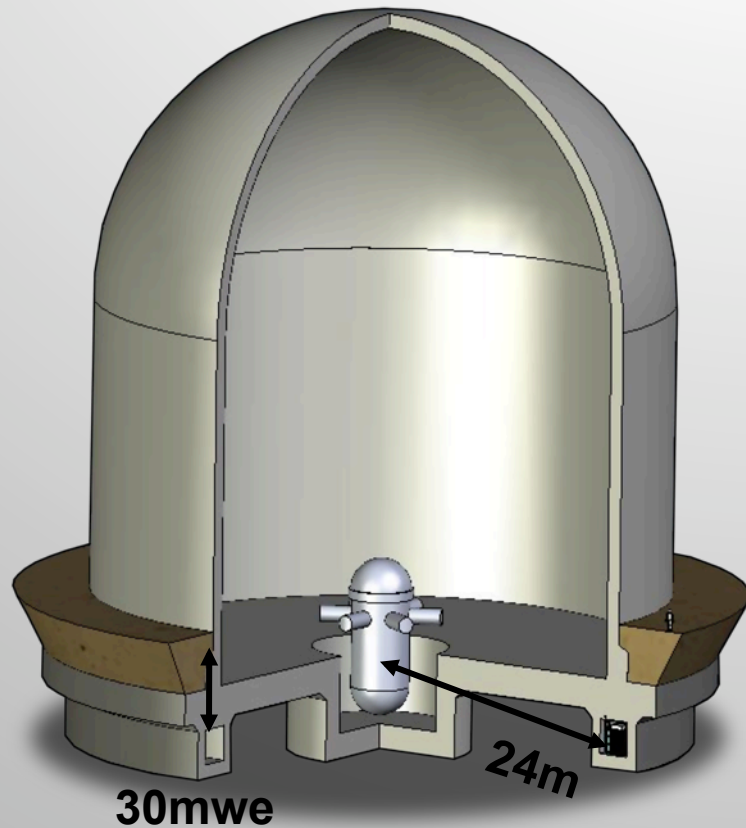


Direct Observation of reactor fuel burnup via antineutrino counting



- We have cultivated an exceptionally strong and trusting relationship with SONGS:
  - A multitude of access requests have been readily granted since 1999
  - Provide unescorted reactor access, deployment assistance, commercially sensitive fueling data, introductions to other operators, .....
- We possess unparalleled operational experience in this industrial environment:
  - **Five** detector deployments since 2003

# Tendon Galleries are Ideal Deployment Locations



- High Flux:  $\sim 10^{17}$  n/m<sup>2</sup>/s
- 130-180m to other reactor
- Gallery is annular – unfortunately no possibility to vary baseline





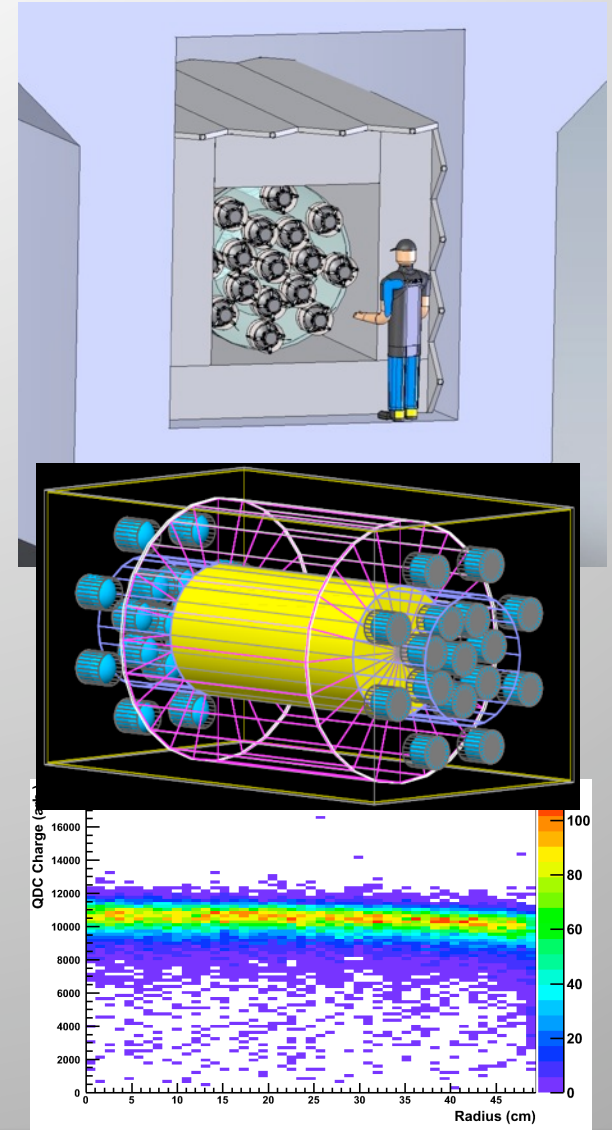
# Tendon Gallery Access





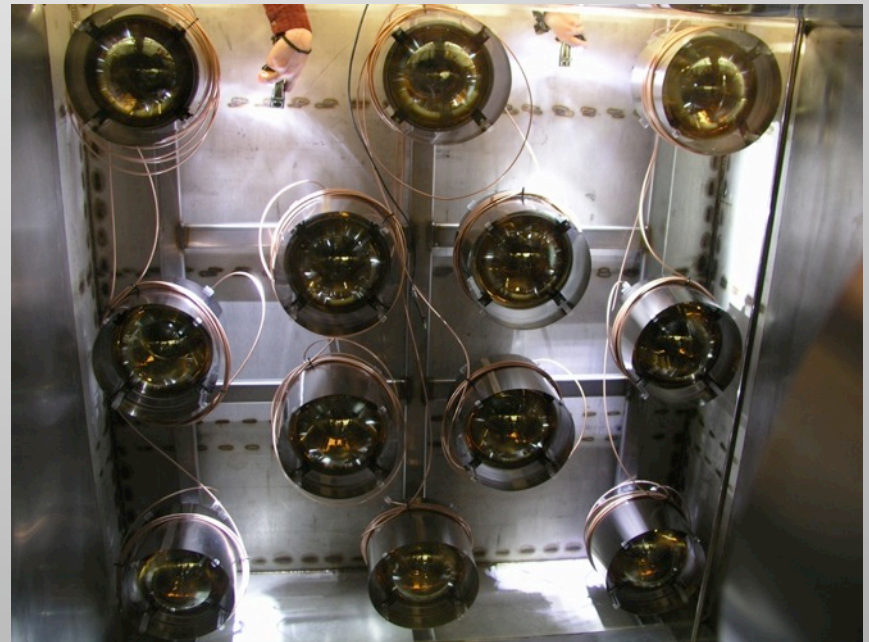
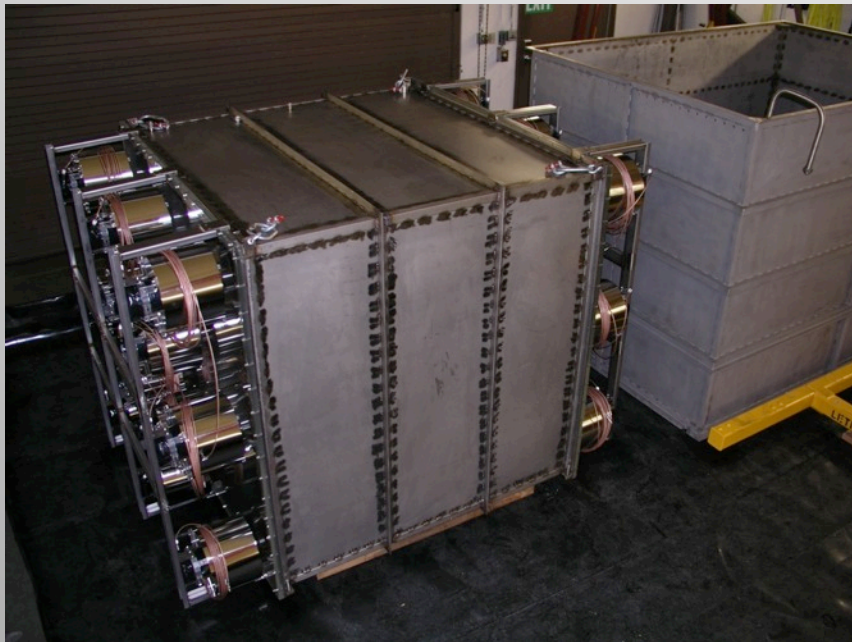
# The SCRAAM Detector Concept

- A relatively long/narrow geometry is needed:  $\text{Ø}1\text{m} \times 2\text{m}$  length
  - Tendon gallery is fairly narrow
- 1.5 ton active mass
  - ~9000 inverse beta interactions/day
  - Conservative 40% efficiency gives detection rate of ~4000/day
- Double ended optical readout and diffuse reflective coating for good light collection and position uniformity: expect to achieve 10% energy resolution at 1MeV
- Guide tubes for calibration
- Aim for at least 4% absolute normalization
  - e.g. include partial “gamma catcher” to increase precision and efficiency
- Component costs: ~\$800k



# We have completed considerable R&D on detectors of this scale

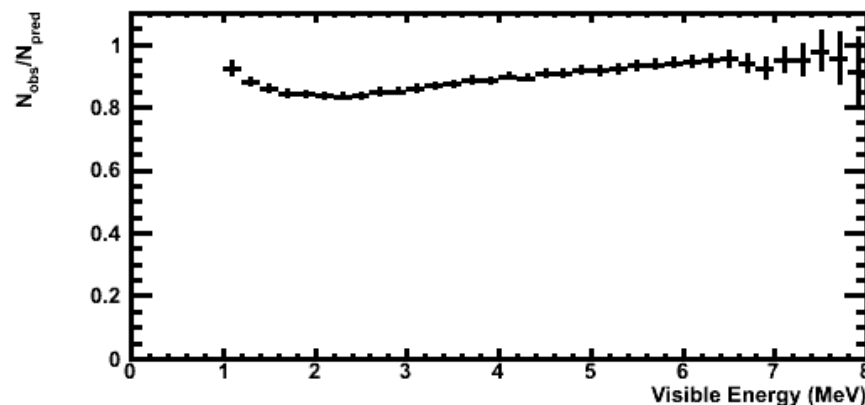
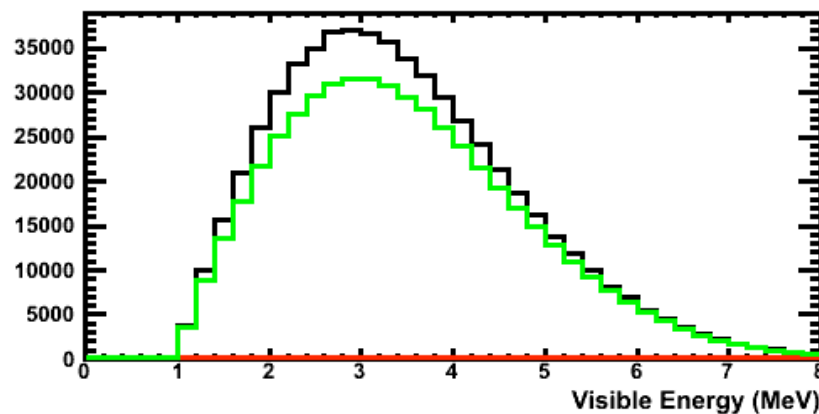
- Most recent: 3.6 ton liquid scintillator detector (BC-525, 0.1% Gd)
  - For deployment at a CANDU6 reactor in 2012
  - Understand safety and regulatory requirements for reactor site
  - Successful commissioning run just completed
  - Validated mechanical design for double ended PMT readout



# Example Oscillation Patterns: For SONGS core, spectral sensitivity remains at 24m

150 days,  $\sin^2(2\theta) = 0.165$ ,  $\Delta m^2 = 0.15 \text{ eV}^2$

1.5% bin-to-bin systematic, 8/1 Signal/Background

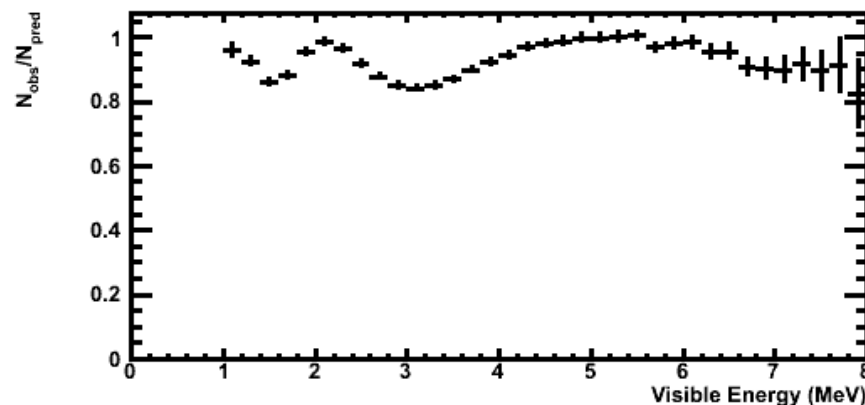
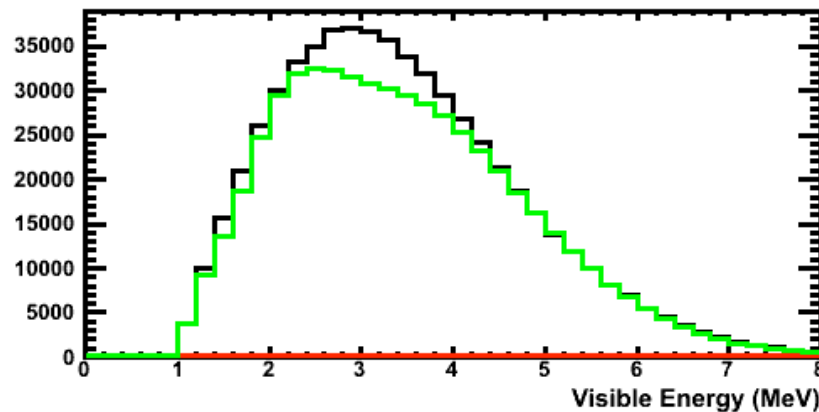




# Example Oscillation Patterns: For SONGS core, spectral sensitivity remains at 24m

150 days,  $\sin^2(2\theta) = 0.165$ ,  $\Delta m^2 = 0.60 \text{ eV}^2$

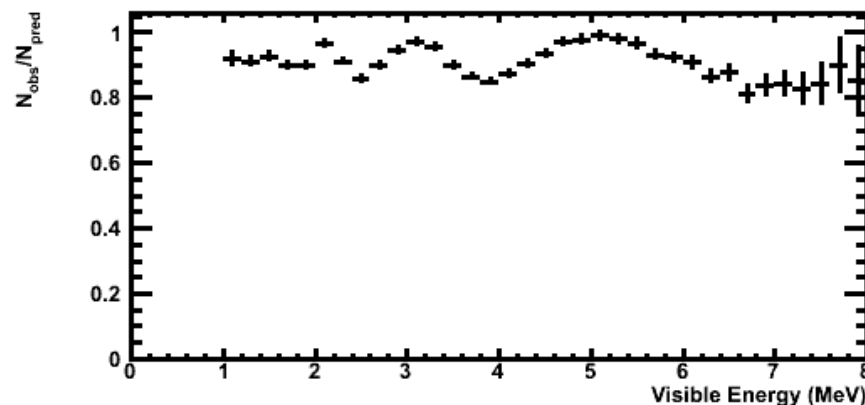
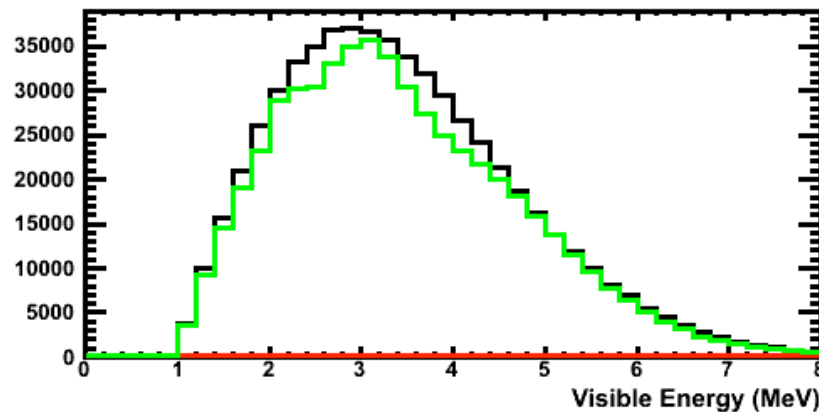
1.5% bin-to-bin systematic, 8/1 Signal/Background



# Example Oscillation Patterns: For SONGS core, spectral sensitivity remains at 24m

150 days,  $\sin^2(2\theta) = 0.165$ ,  $\Delta m^2 = 1.2 \text{ eV}^2$

1.5% bin-to-bin systematic, 8/1 Signal/Background

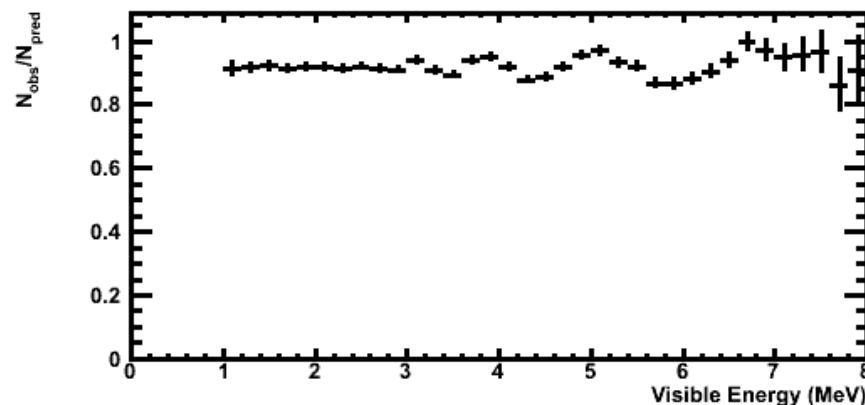
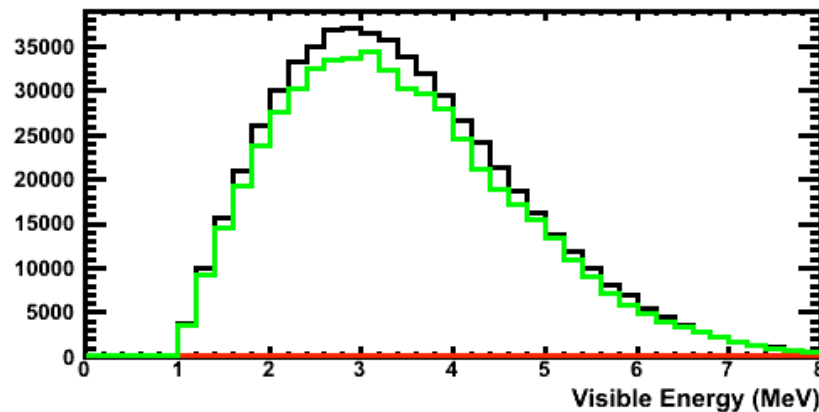


# Example Oscillation Patterns:

## For SONGS core, spectral sensitivity remains at 24m

150 days,  $\sin^2(2\theta) = 0.165$ ,  $\Delta m^2 = 2.4 \text{ eV}^2$

1.5% bin-to-bin systematic, 8/1 Signal/Background



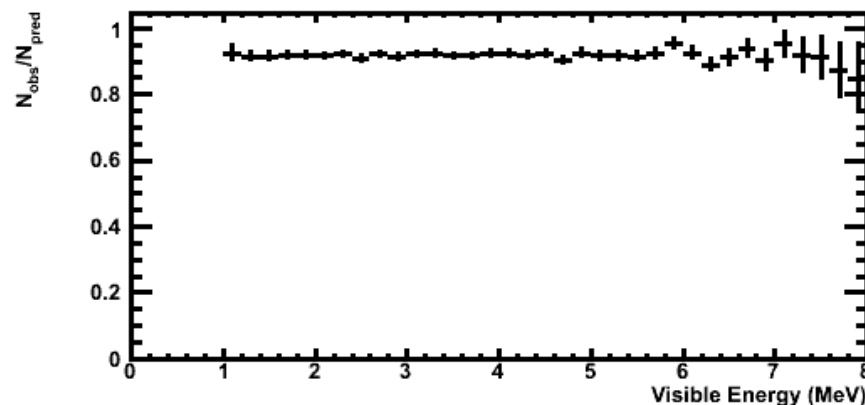
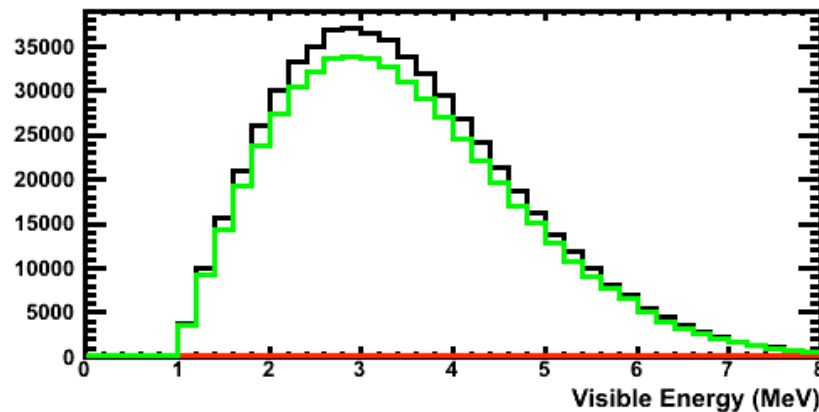


# Example Oscillation Patterns:

## For SONGS core, spectral sensitivity remains at 24m

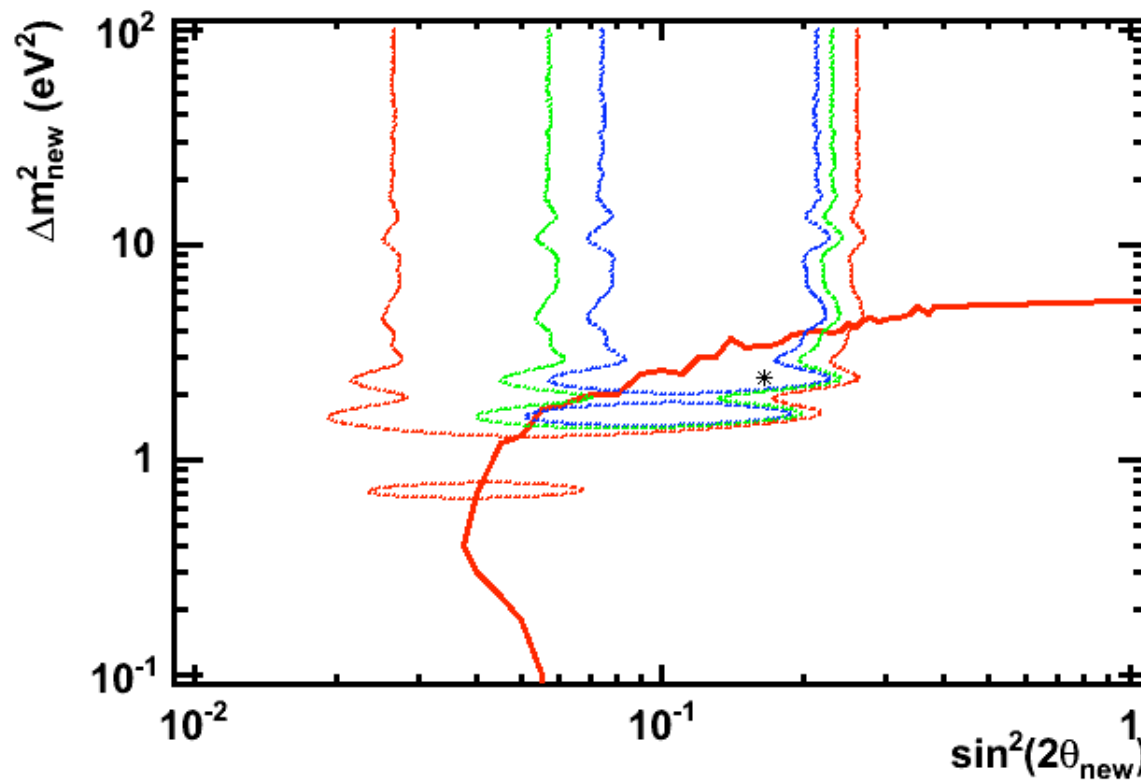
150 days,  $\sin^2(2\theta) = 0.165$ ,  $\Delta m^2 = 4.8 \text{eV}^2$

1.5% bin-to-bin systematic, 8/1 Signal/Background



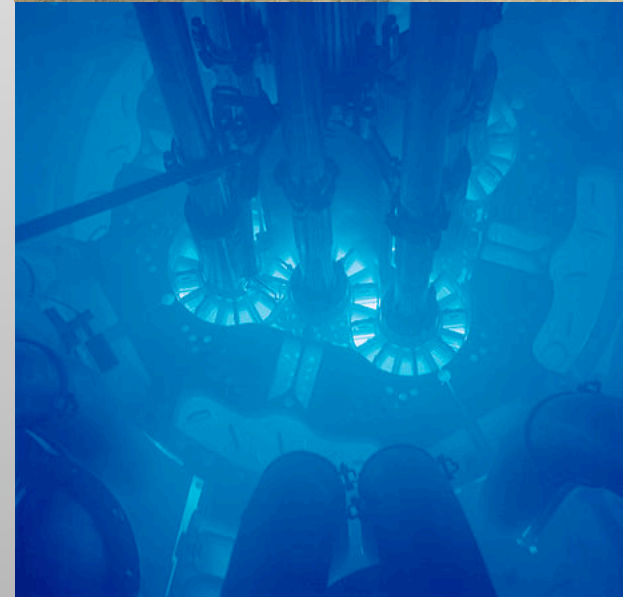
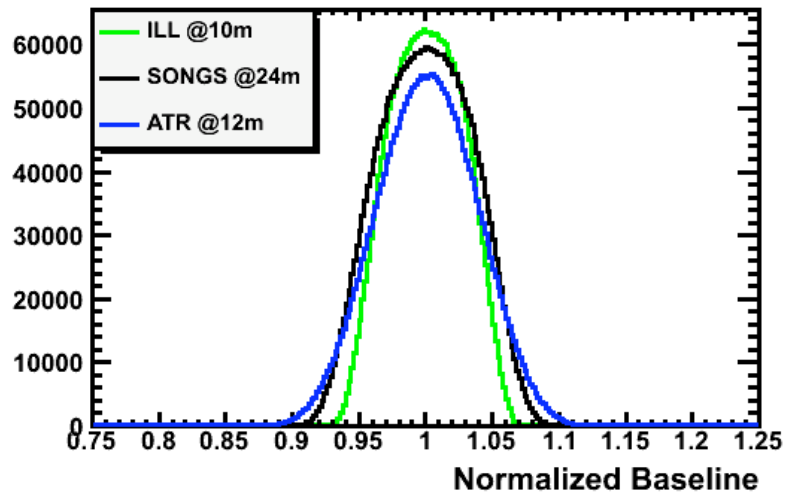
# Exclusion Estimates: Shape

- 150 days, 99% C.L.
- 1.5% Energy scale error, 8/1 Signal/Background



# Investigating possibility of complementary compact core measurement

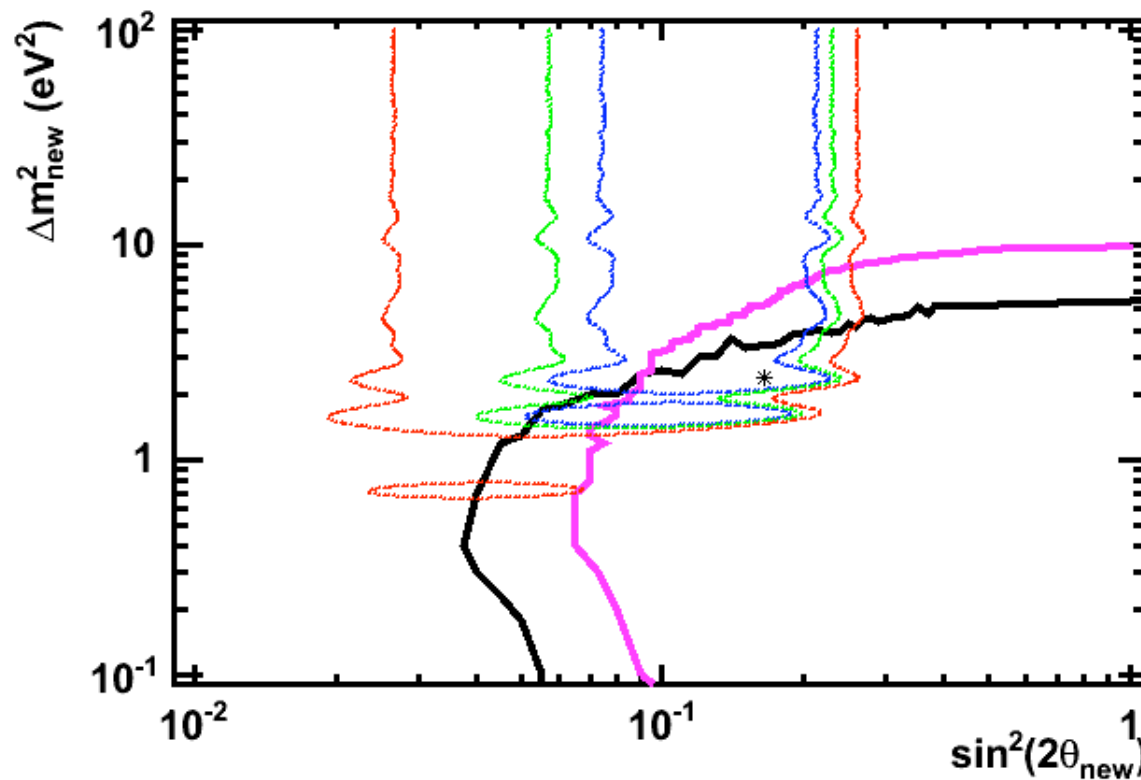
- Advanced Test Reactor at Idaho National Lab
- Unique “serpentine” 1.2m HEU core,  $\sim 150\text{MW}_{\text{th}}$
- Convenient 60 day on, 30 day off cycle
- Potential below grade deployment locations near core
- At 12m baseline, spread similar to that at SONGS





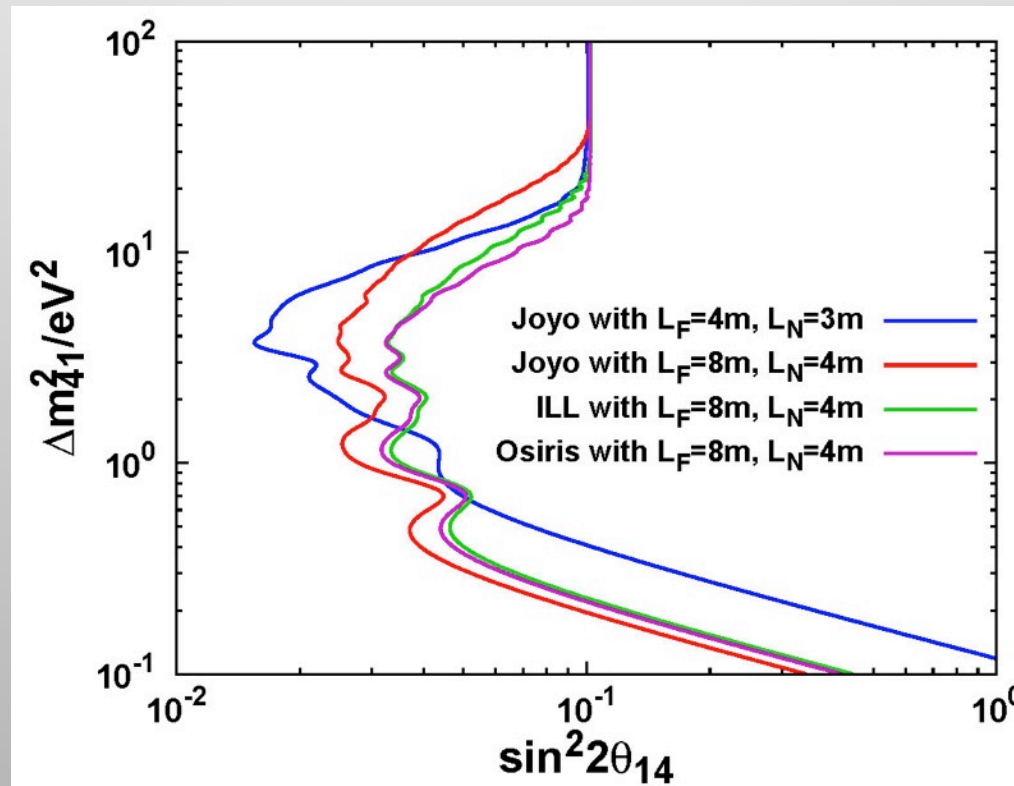
# Exclusion Estimates: Shape

- 99% C.L.; 150 days@ SONGS; 300 days @ ATR
- 1.5% Energy scale error, 8/1 Signal/Background



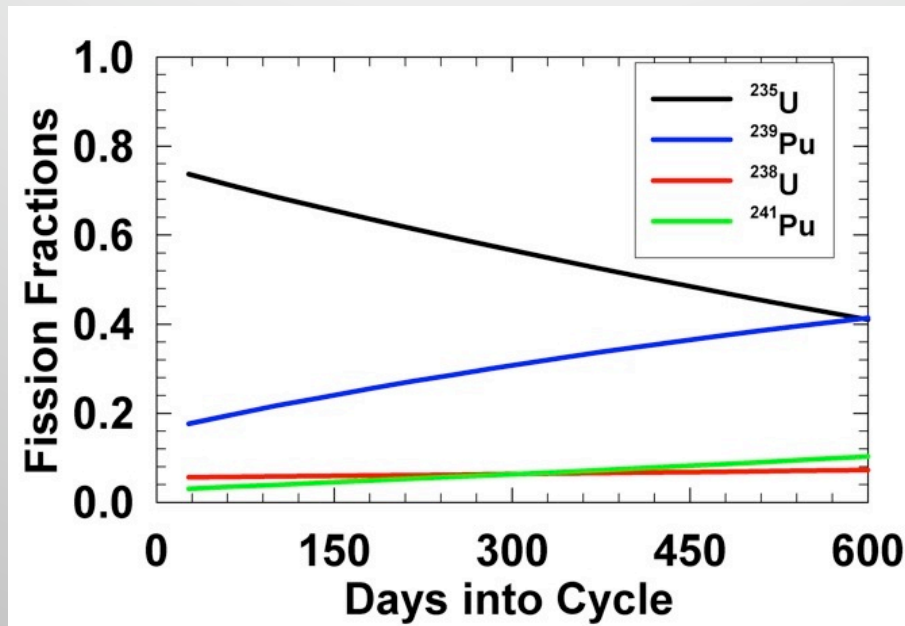
# Combined analysis of multiple baselines would have broad sensitivity

- Key parameters would be relative energy scale and normalization: O. Yasuda, arXiv:1107.4766



# SONGS Core evolution is well understood

- Again, through our long interaction with SONGS we have access to operator fueling and reactor data

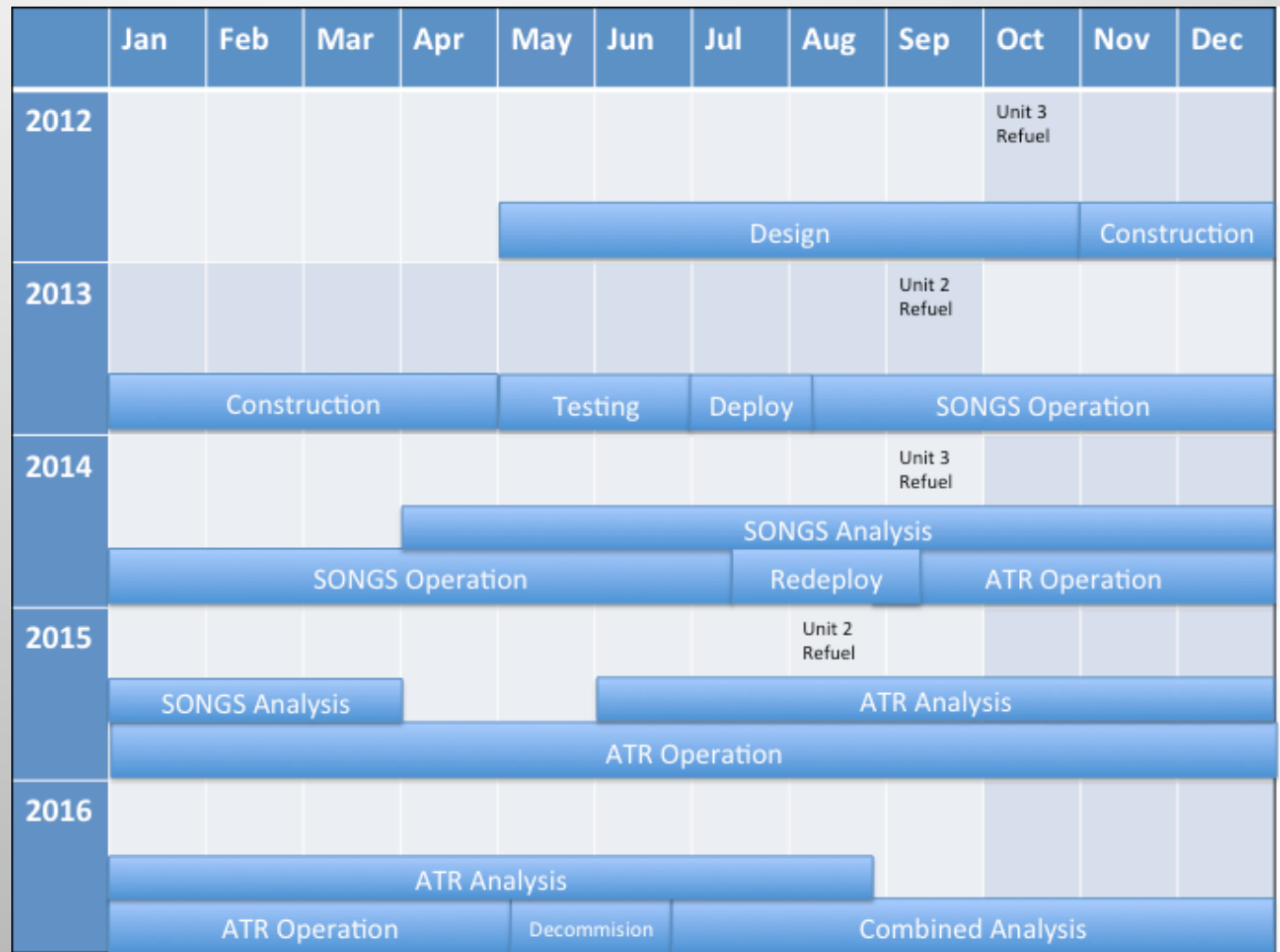


- Unlike the theta13 near detectors, the SCRAAM spectrum measurement would effectively be from a single core
  - In the absence of spectral distortion, this measurement could better constrain prediction uncertainties



# Nominal Schedule

- SONGS outages are key; ~50 day background measurement:
  - Unit 2 Sept. '13
  - Unit 3 Sept '14
- Given our recent experience, 15-18 months from design to deployment seems feasible
- Could have first results within ~9 months of data taking



# Effort and Budget

	FY12	FY13	FY14	FY15	FY16
PI	0.5 FTE	0.5 FTE	0.75 FTE	0.75 FTE	0.75 FTE
Post-doc	0.5 FTE	0.5 FTE	0.75 FTE	0.75 FTE	0.75 FTE
Engineer/ Technologist	0.66 FTE	0.33 FTE	0.17 FTE	0.17 FTE	0.08 FTE
Total FTE	1.66 FTE	1.33 FTE	1.67 FTE	1.67 FTE	1.58 FTE
Equipment	\$350K	\$440K			
Approx. Total Cost	\$900K	\$900K	\$500K	\$500K	\$500K

# Potential Collaborators

- John Learned, U. Hawaii
  - Fabrication, software
- HANARO Group (South Korea)
  - High flash-point Scintillator
- ...





# Conclusions

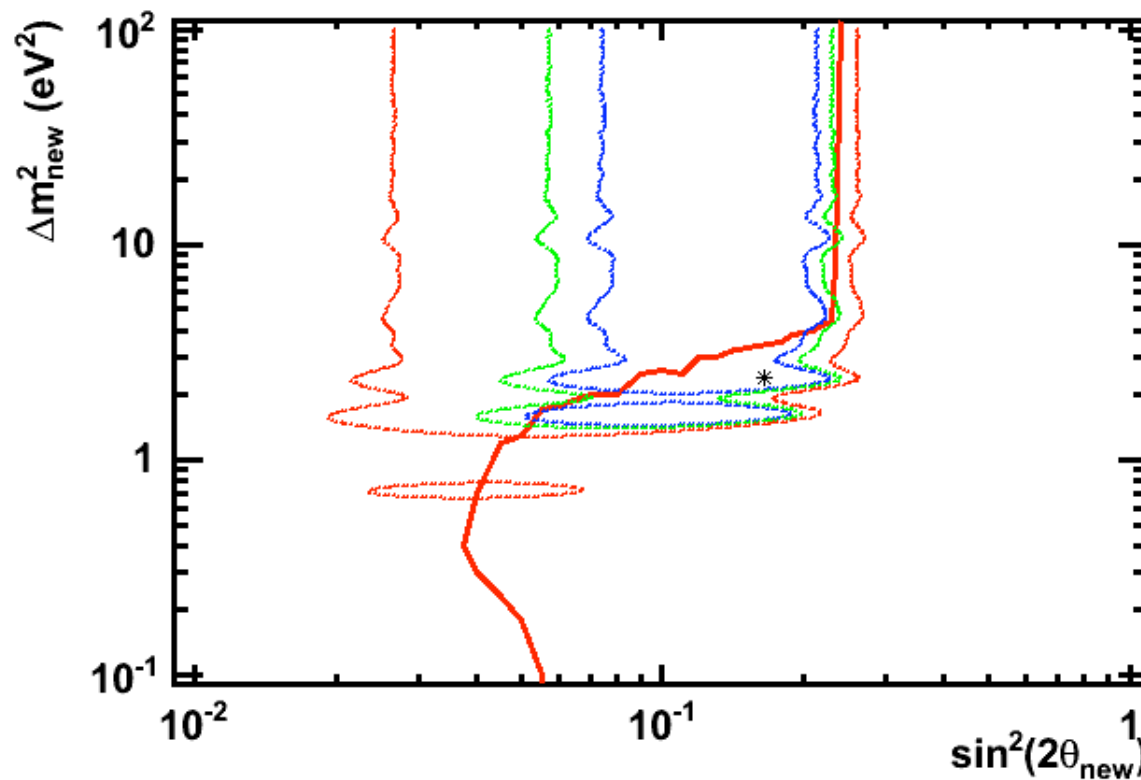
- Short baseline reactor efforts have continued, attempting to develop a new safeguards technique
  - The reactor access, reactor simulation, and detector design expertise from the applied community can be exploited to probe the “RAA”
- Short baseline measurements at appropriate small (research) and large (power) reactors would be complementary:
  - Efficiently probe different  $\Delta m^2$  regions and measuring spectra from different core compositions
  - SONGS appears optimal for a power reactor deployment
  - ATR appears very promising as a research reactor deployment site
  - Combined analysis of two deployments could have even better sensitivity
- SCRAAM would rapidly exclude a large fraction of the  $\sim 1\text{eV}^2$  “RAA” allowed phase space, and have good discovery potential in the “best-fit” region





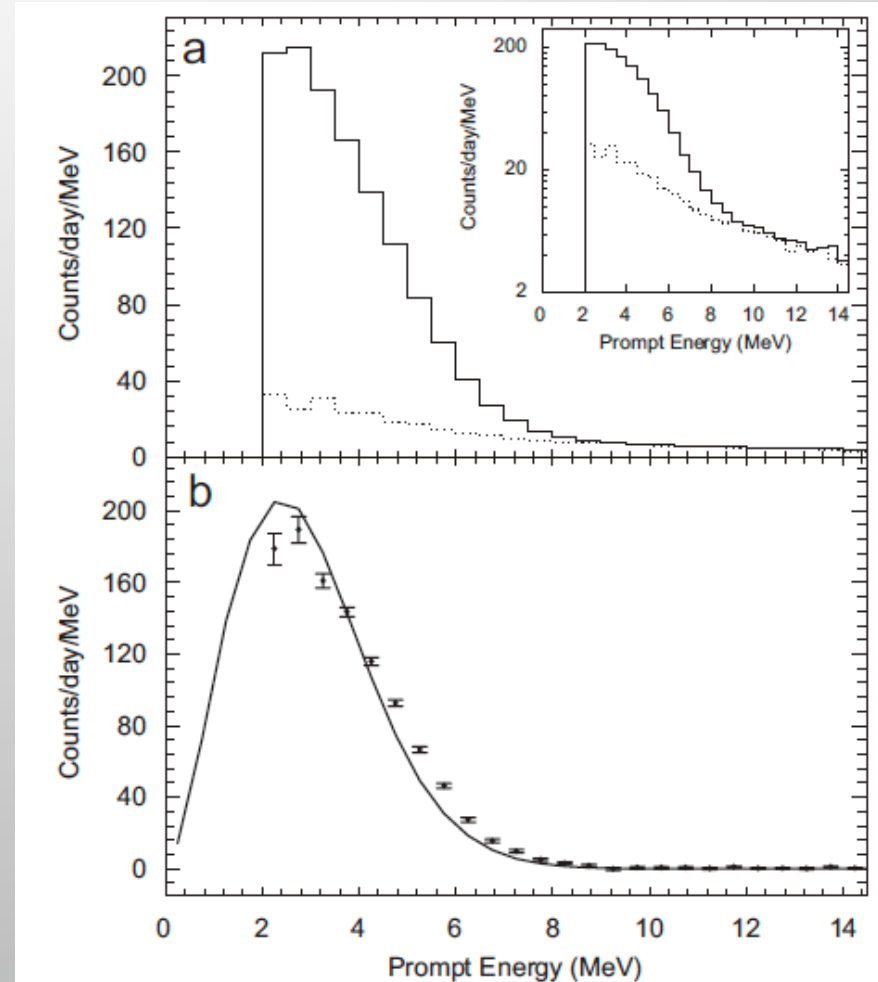
# Exclusion Estimates: Shape + Rate

- 150 days, 99% C.L.
- 4% Normalization, 1.5% Energy scale error, 8/1 Signal/Background



# SONGS Backgrounds

- Our SONGS<sub>1</sub> detector had S/B of  $\sim 4/1$
- Background was primarily:
  - Fast neutron recoil followed by capture
  - Multiple neutron capture
- There is reason to believe that we can do considerably better with SCRAAM:
  - SONGS<sub>1</sub> had only 95% muon veto and “non-hermetic” shielding
  - Improved neutron capture efficiency and analysis will allow rejection many more multiple neutrons





# Detection Rate, Detector Systematics

- Assuming 40% efficiency, expect detection rate of about 4000 v/day
- Precision on absolute efficiency of  $\sim 4\%$  would require considerable effort, but appears feasible
- Extensive source calibrations would be required

Systematic	Target Value	Mitigation
No. of target protons	1.5%	Weighting, solvent selection
Neutron efficiency	1.5%	Gamma catcher, calibration
Positron efficiency	1%	calibration, ideally $\approx 500$ keV threshold
Core-Detector distance	0.5%	Through document review, possible survey
Deadtime	0.25%	precise measurement, tracking
Detector Total (Flux Measurement)	2.4%	
Reactor Systematics	2.7%	
Total on $N_{obs}/N_{pred}$	3.6%	



# A compact core effort: Nucifer

(see also Y.D. Kim poster)



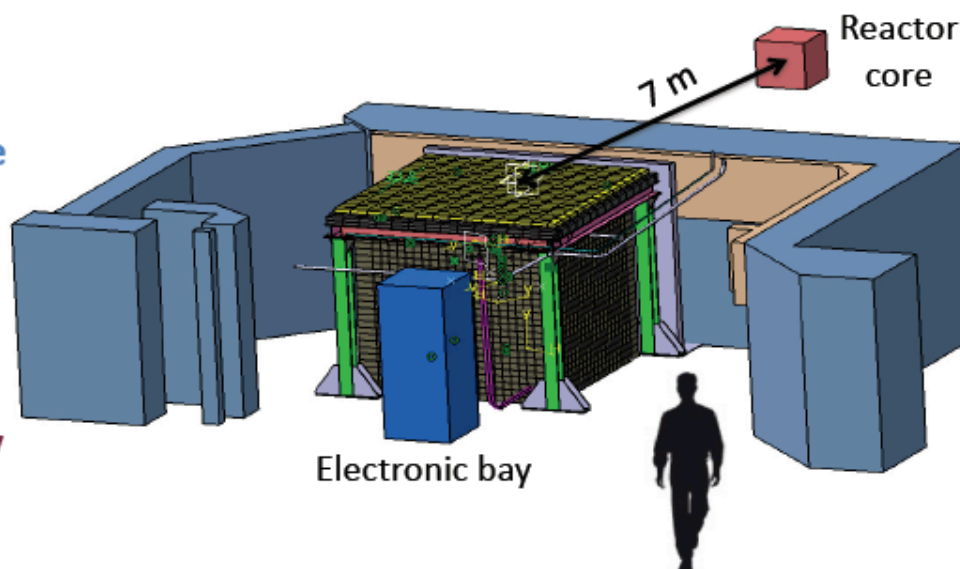
Nucifer

## Nucifer @ Osiris



- 70 MW reactor
- Nucifer 7 m from the core
- 15 mwe overburden

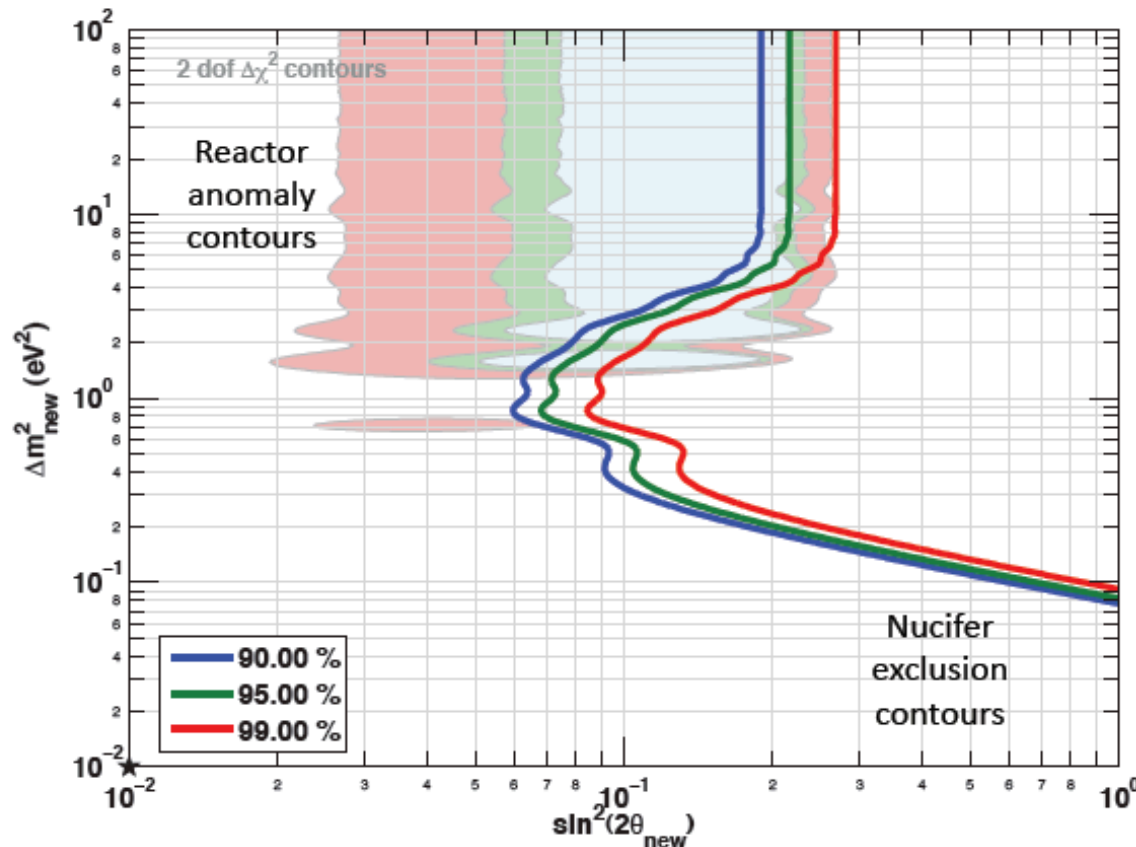
650  $\nu$ /day expected  
Assuming 50% det efficiency



- **Reactor Background:**
  - Additional 10 cm lead shielding needed due to reactor induced  $\gamma$  rays
- **Based on simulation and on site measurements:**
  - $S/B_{\text{accidentals}} = 1$
  - $S/B_{\text{correlated}} = 0.25$  before PSD cut,  $\sim 2.5$  expected after PSD selection.  
Reactor OFF 33% of the time, will allow final background subtraction.



# Testing the 4<sup>th</sup> $\nu$ hypothesis



**100 days full power  
@ Osiris:**

- 4% norm error
- E resol = 0.15\*E
- 2% E scale error
- S/B = 1 (?),  
assuming same  
shapes (worst case).

# Other Proposed Efforts (\$\$\$)

- PBq neutrino sources into KamLAND, SNO+, etc
- Multi-Detector Accelerator experiments
  - BOONE
  - 2 x LAr TPC @ CERN PS
  - 2 x LAr TPC @ FermiLAB
- (PLANCK will provide much tighter cosmological bounds in ~2014-15)



